

# Observation of Dust in DIII-D During Normal Plasma Operation Using Rayleigh/Mie Scattering

by  
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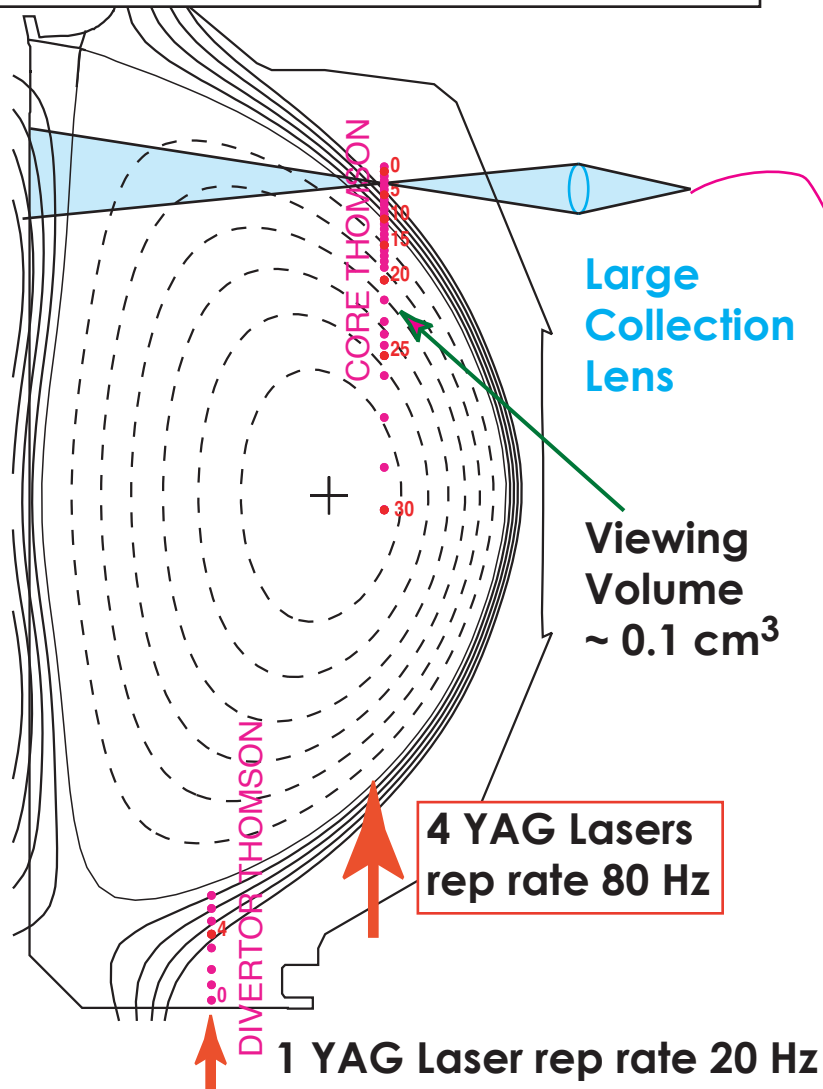
**Feb. 28, 2006**

# Multiple Sources of Dust Exist in DIII-D

- **Flakes from redeposited  $C_xD_y$  films**
- **High Heat Flux Generation**
  - Leading edges and monopolar arcs
  - Thermal Stress Induced Fracture
  - Volume Condensation
- **Particles left over from entry vent activities**
- **Degradation of grafoil compliant layers under tiles**

# DIII-D Thomson System Used to Detect Dust Particles

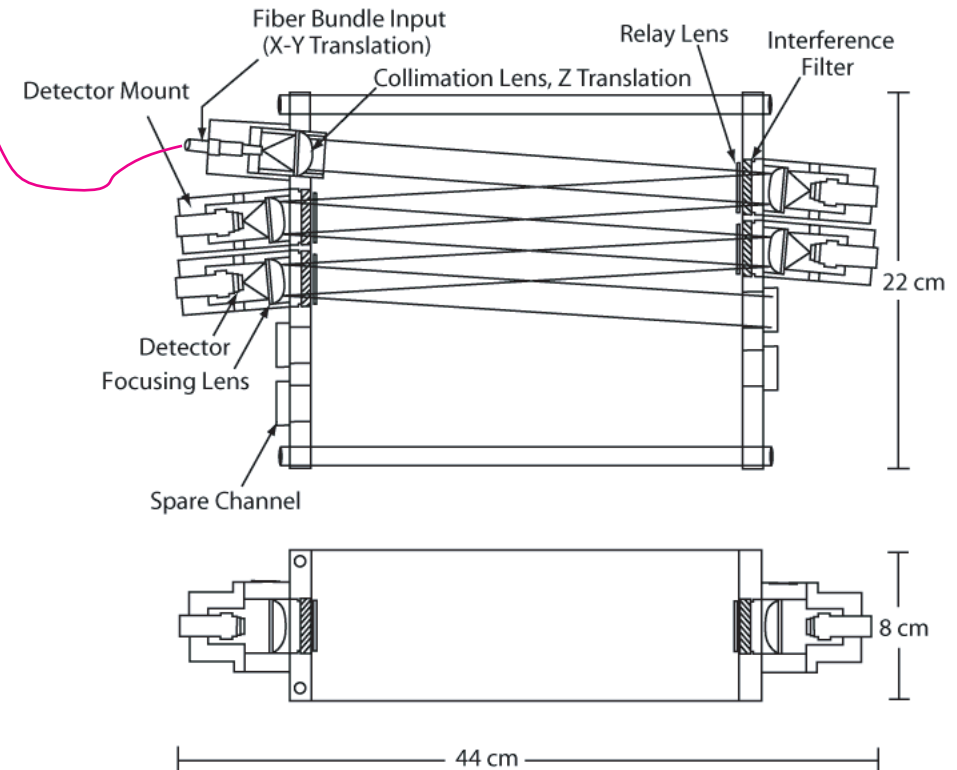
## • Thomson viewing locations



## • Polychromator at each location

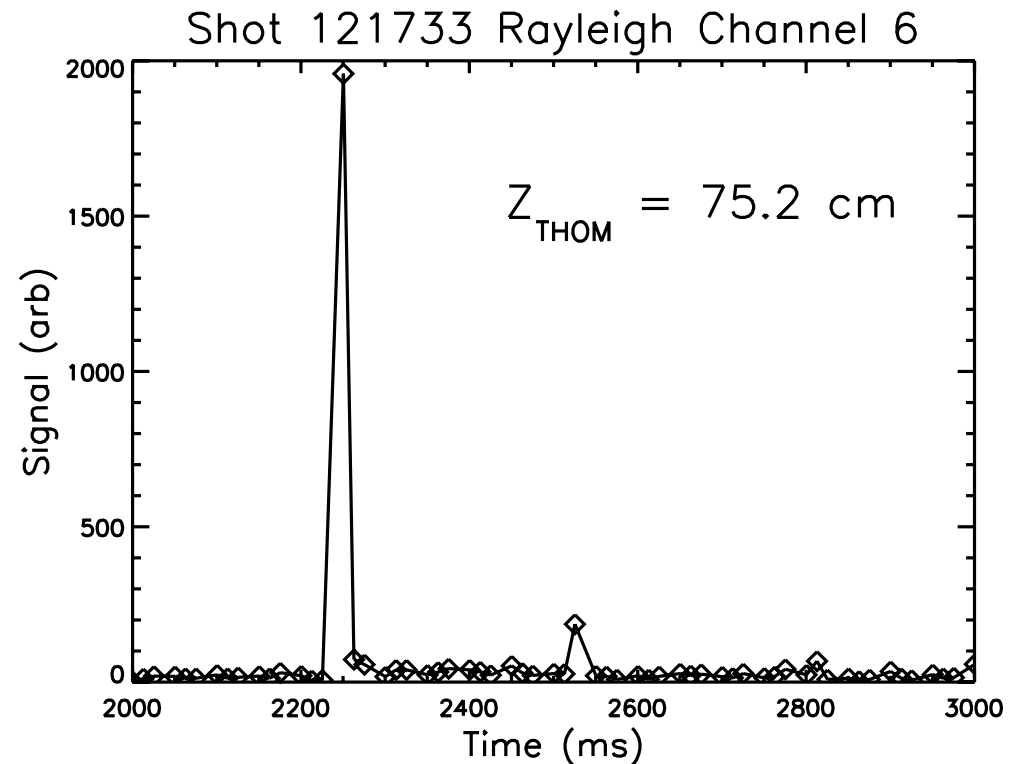
- Rayleigh Channel Passes Laser Light at 1064 nm
- Next Channel at 1062 nm has  $10^{-2}$ - $10^{-4}$  extinction at laser line
- 4 more channels further blue shifted

Fiber Optics



# Large Excursions Observed in Rayleigh Signal

- **Other causes of large excursions can be discriminated against.**
  - Bright Thermal/plasma light from far wall
    - Discriminated against by fast background subtraction
    - Would appear simultaneously in several spatial and  $\lambda$  channels
  - Scattered laser light from wall
    - Would appear simultaneously in several spatial channels
  - Neutrons,  $\gamma$ 's, cosmic rays
    - Would appear uniformly in all channels



- **Observed during normal plasma operation**

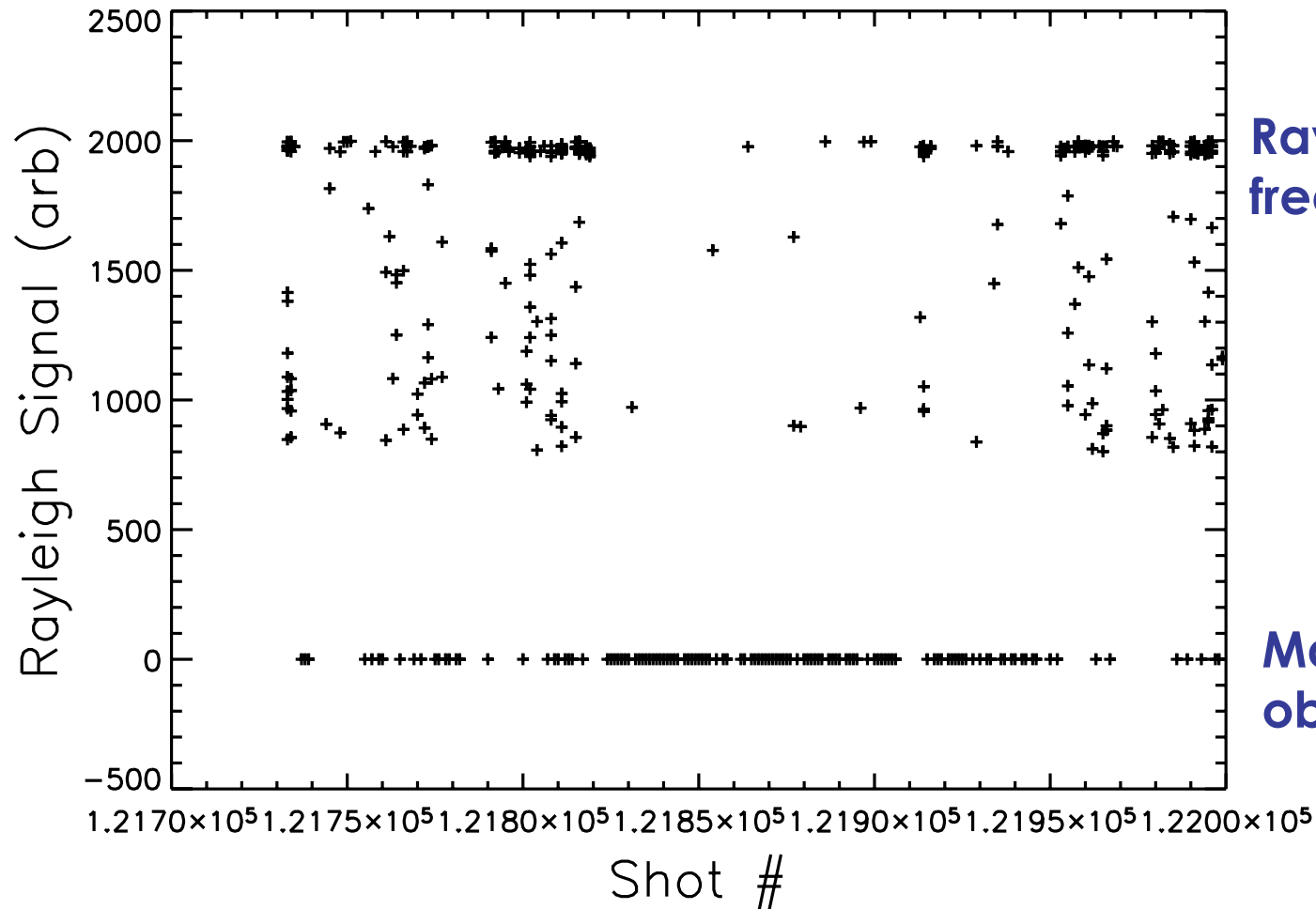
# Dust signatures found in search of existing Thomson data

- **Examined core and divertor data over ~700 discharges**
  - Search of 31 core channels:  $1.5 \times 10^5$  laser pulses
  - 8 Divertor channels:  $3.8 \times 10^4$  laser pulses
- **Search only when  $I_p > 0.9$  MA, for at least 700 ms**
  - Require Rayleigh signal  $> 200$  counts
  - Reject times with simultaneous excursion in more than 2 core (divertor) Thomson spatial channels
  - Reject times when neighboring times have large excursion in same spatial channel
  - Excursions are recorded for all  $\lambda$  channels

# Search Statistics From Core Thomson shots 121733-121999

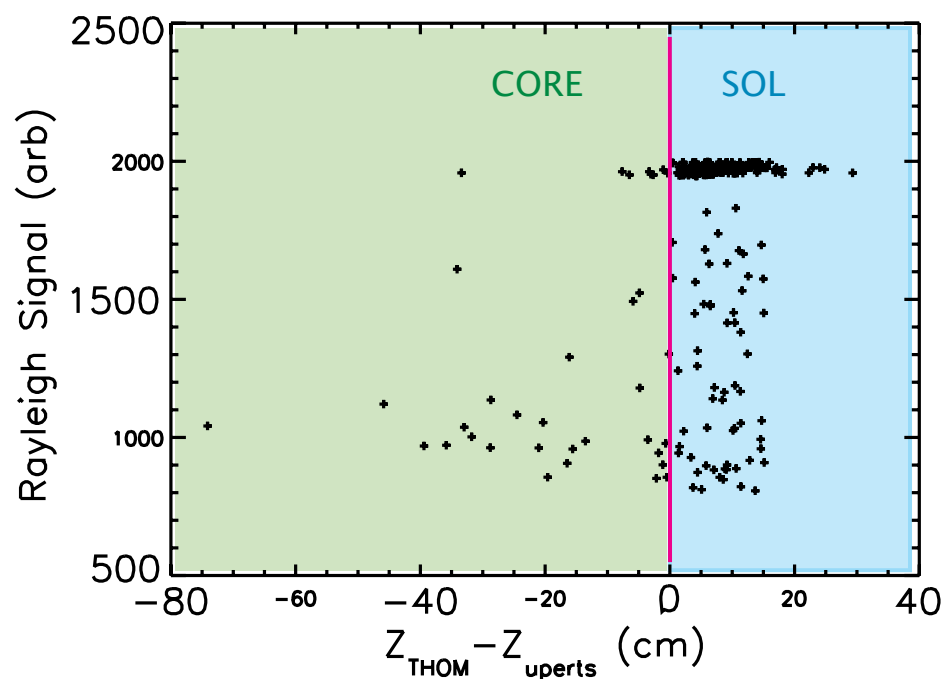
- **209 Shots with at least 700 ms at  $I_p > 900$  kA.**
  - 124 Shots exhibited no dust signature
- **293 large events in Rayleigh channels selected**
  - 266 events in region of typical SOL (13 channels)
  - 27 events in region of confined plasma (18 channels)
  - 28 events in shifted  $\lambda$  channels (150 channels)
  - 26 times rejected due to excursions in multiple Rayleigh channels

# Dust apparant in recent shots taken over a two week period

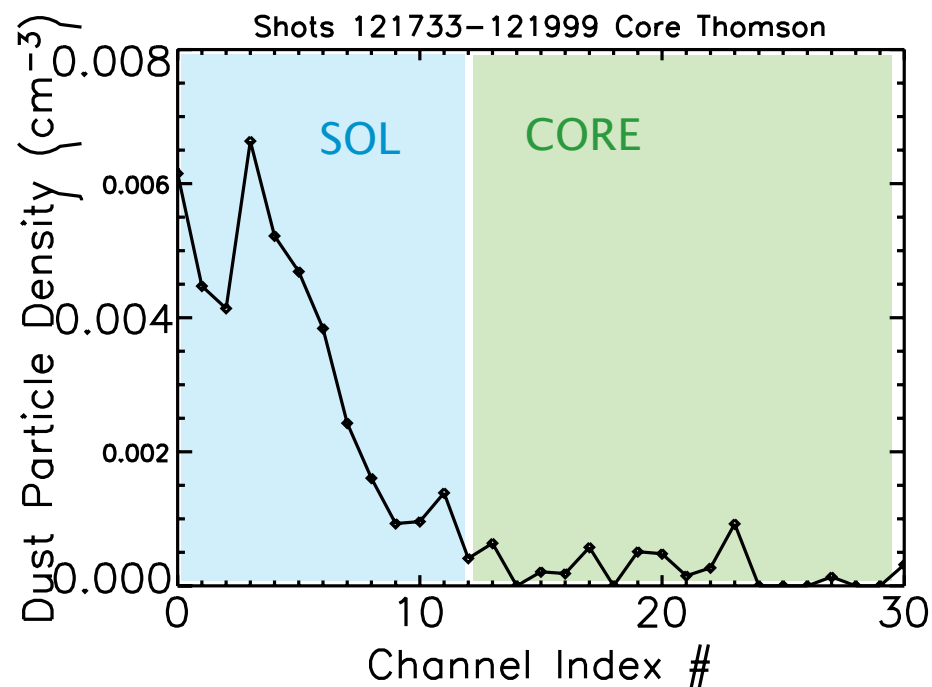


# Dust Density Falls from Wall to Separatrix

Dust Frequent in SOL but  
Rare in the Core



Particle Density vs  
Thomson Channel Index

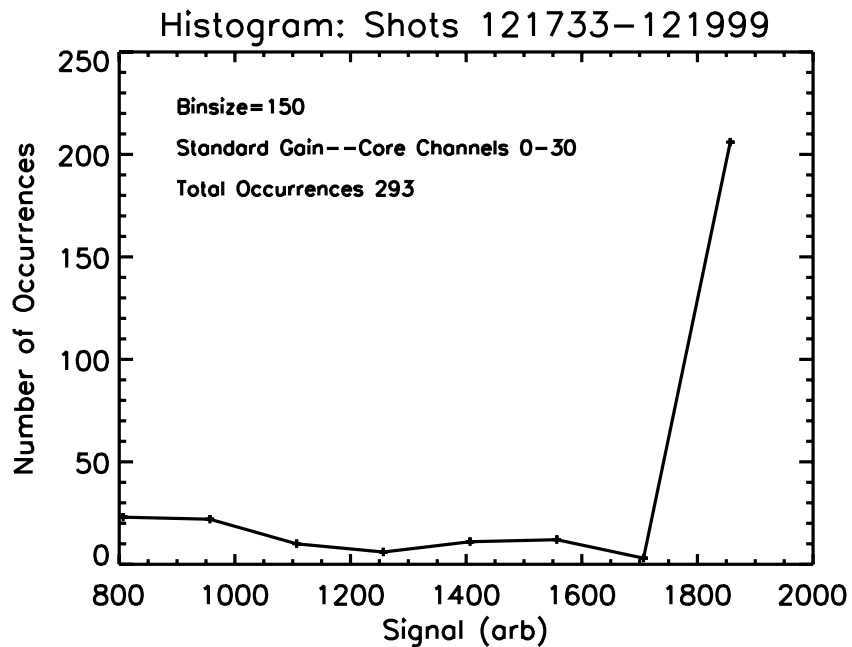


Time Averaged Dust density =  
 $\# \text{ events} / (\text{viewing volume} \times \# \text{ laser pulses})$



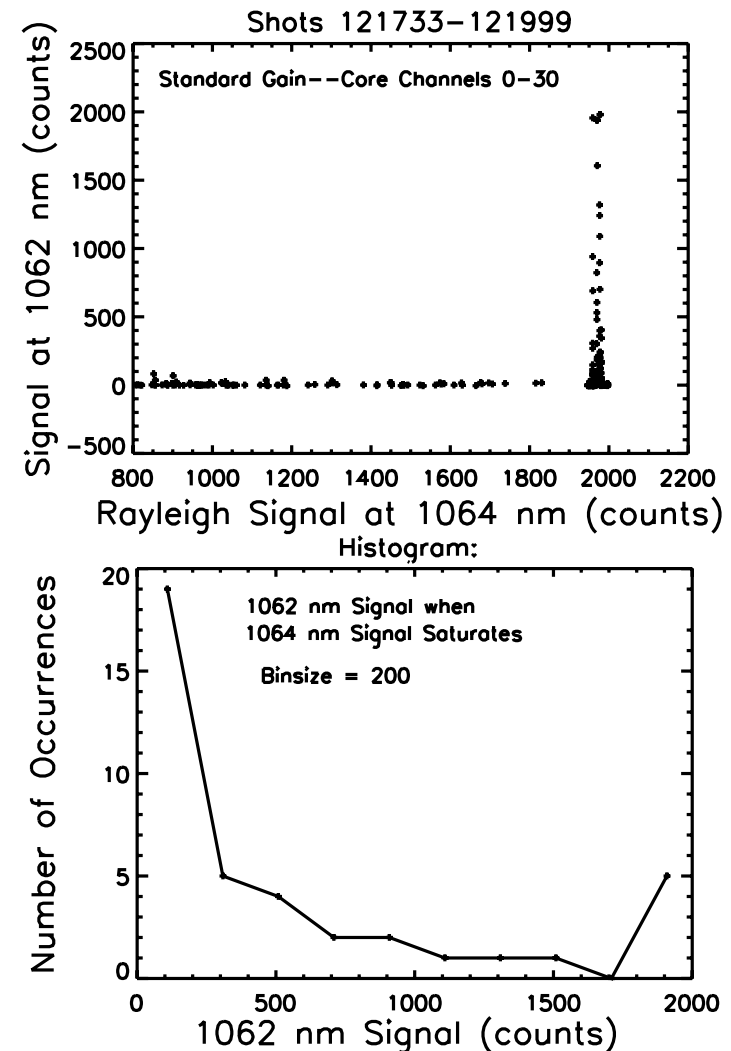
# Rayleigh Signal Typically Saturated

## Rayleigh Channel 1064 nm



- 1062 Channel Attenuates YAG Laser Light by  $10^{-4}$

## 1062 nm Channel



# Rough Estimate of Particle Size from Thomson Rayleigh Calibration

- Rayleigh Scattering Cross Section,  $\sigma$  for conducting particle, radius  $a$

$$\sigma = 10 \frac{\pi}{3} k^4 a^6$$

$$k = 2 \pi / \lambda$$

$$(a_{\text{part}}/a_{\text{Argon}})^6 = \frac{S_{\text{part}}}{S_{\text{Argon}}/N_{\text{Argon}}}$$

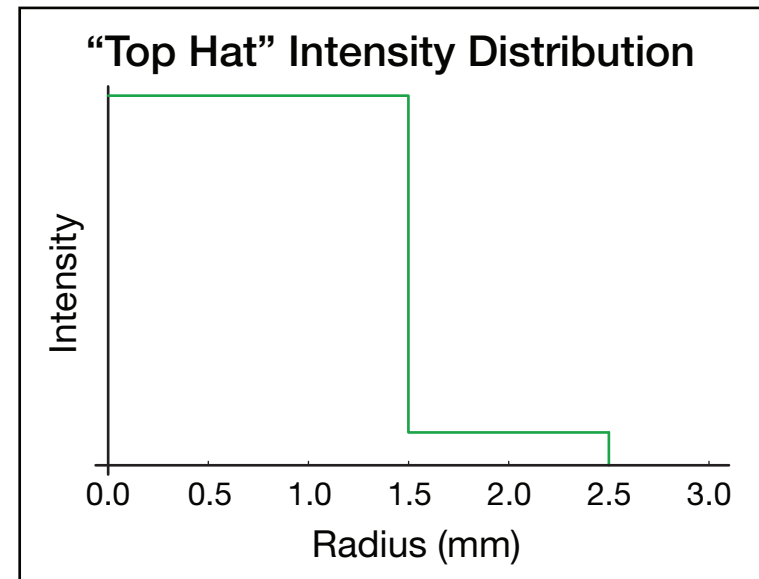
- $S_{\text{Argon}}$  = Rayleigh Signal with  $N_{\text{Argon}}$  atoms in viewing volume  
–  $a_{\text{Argon}}$  ,  $\sim 1.1 \text{ \AA}$ , obtained from measured  $\sigma_{\text{Argon}}$  from the literature

# Modeling goals

- **Laser intensity is nonuniform over imaging volume, denying the existence of a one-to-one function from particle radius to detected signal size**
  - We can only take a particle radius [distribution] and predict a signal size [distribution]
- **As such, our method involves:**
  - Using an assumed particle radius distribution to predict a signal size distribution;
  - Calculating the fit error between the predicted and experimental signal size distributions; and
  - Adjusting parameters in the assumed particle radius distribution to minimize error in the predicted signal size distribution

# Model assumptions

- **Rayleigh scattering formula holds**
  - In reality, only a good approximation if scattering particle radius is less than about 10% of wavelength of incident light
  - $a < 106 \text{ nm}$ , in this case
- **“Top hat” laser intensity distribution**
  - Our data suggests:
    - 95% of laser power within 1.5 mm radius
    - 5% in concentric ring of 2.5 mm outer radius
  - We use this “top hat” approximation in absence of more precise data from manufacturer



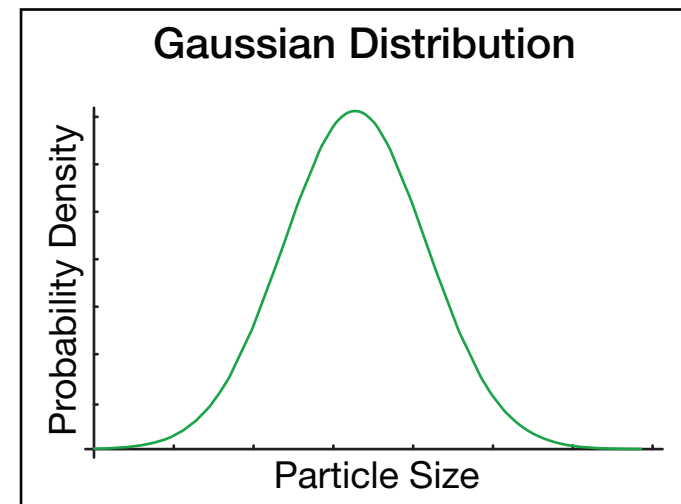
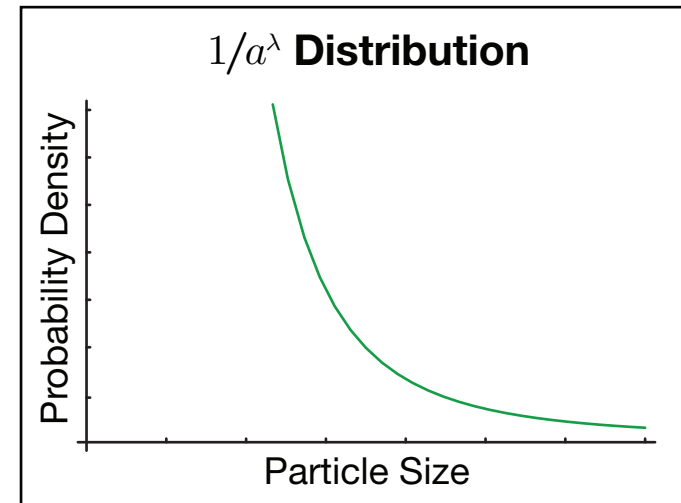
# Particle radii distributions applied

- **Gaussian distribution**

- $f(a) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left(-\frac{(a-\mu)^2}{2\sigma^2}\right)$ 
  - Location parameter  $\mu$
  - Spread parameter  $\sigma$
- Standard distribution for many physical phenomena

- $1/a^\lambda$  **distribution**

- $f(a) = \frac{1}{a^\lambda}$ 
  - Spread parameter  $\lambda$



# From particle radius to signal size

- **Predicting a signal size distribution from a particle radius distribution**
  - We have a particle radius distribution function
  - We know how to find particle radius in terms of signal size (from Rayleigh law)
  - Thus, we simply plug in our formula for particle radius into our particle radius distribution function, and adjust the differential
  - This yields a signal size distribution function

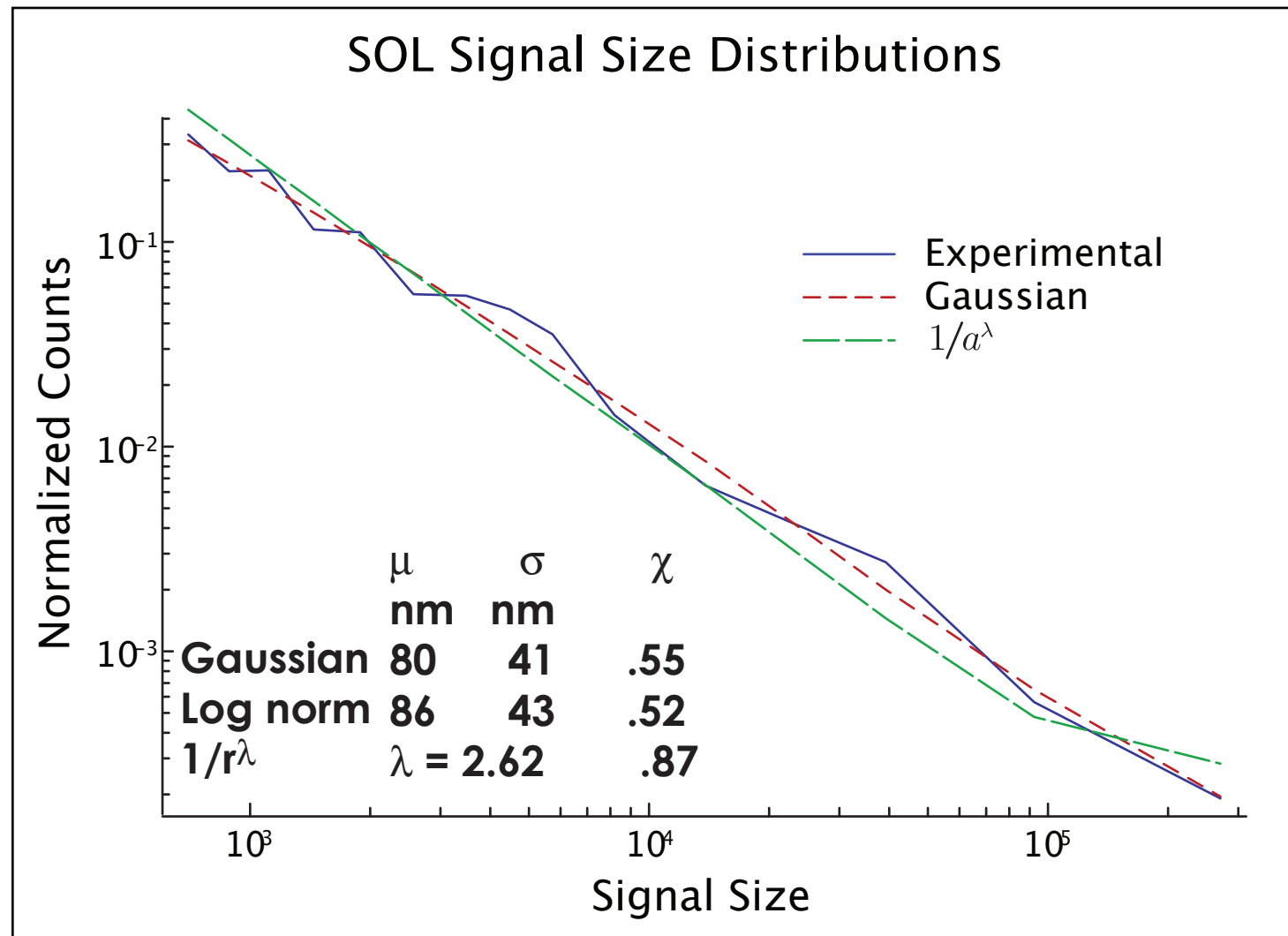
$S$ : signal size,  $\beta$ : calibration factor,  $\Gamma$ : laser photon flux,  $\bar{\Gamma}$ : average laser photon flux

$$f(a) \longrightarrow a(S) = \sqrt[6]{\frac{S \bar{\Gamma}}{\beta \Gamma}} \longrightarrow g(S) = \frac{f(a(S))}{dS/da} = \frac{1}{6S} f\left(\sqrt[6]{\frac{S \bar{\Gamma}}{\beta \Gamma}}\right) \sqrt[6]{\frac{S \bar{\Gamma}}{\beta \Gamma}}$$

$f$ : distribution of particle radius  $a$

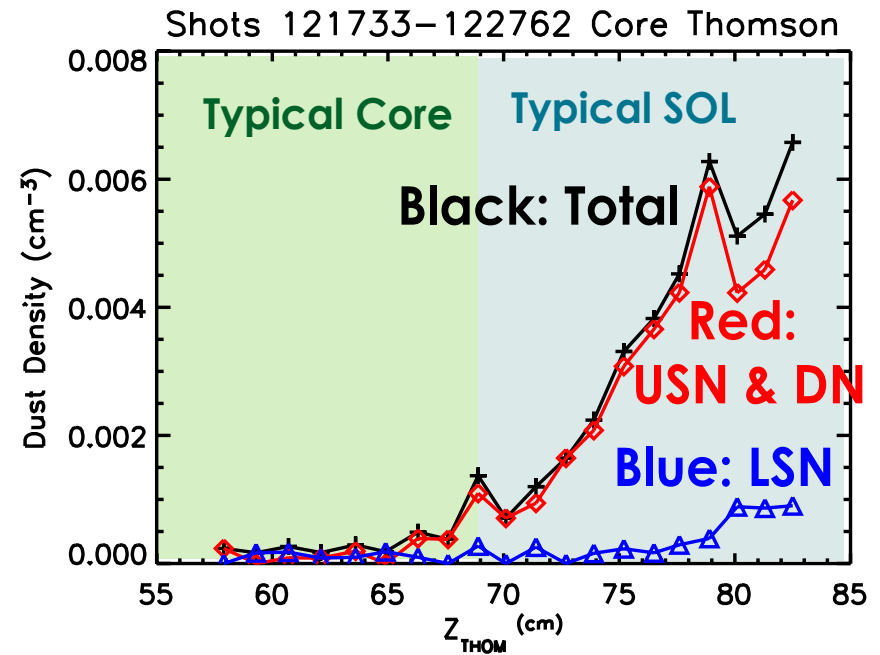
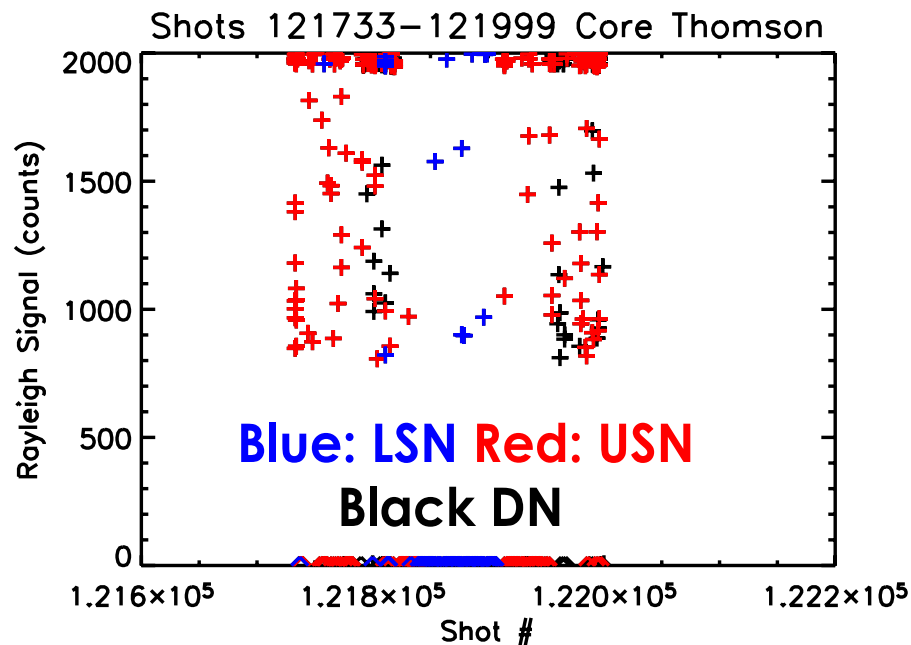
$g$ : distribution of signal size

# Comparison of model distributions to experimental distribution in SOL



# Dust in Upper Region of SOL more Common in Upper Single and Double Null Operation

Dust in Upper SOL clearly more common when upper null is active





# Conclusions

- Dust has been observed in the SOL and divertor regions of DIII-D during normal plasma operations using the existing Thomson scattering system.
- In the SOL the density of dust particles is low,  $\sim 5 \times 10^{-3} \text{ cm}^{-3}$
- The observed pulse height distribution suggests an average particle radius of  $\sim 80 \text{ nm}$ . The existing dataset cannot distinguish between a normal or log normal size distribution
- The equivalent atomic carbon density in the SOL due to the dust is very low ( $\sim 10^{11} \text{ m}^{-3}$ ) and probably does not constitute a significant source of carbon contamination.
- These particles seem to disappear in the SOL and do not reach the core plasma. However if they return to the wall they may constitute a tritium uptake problem.